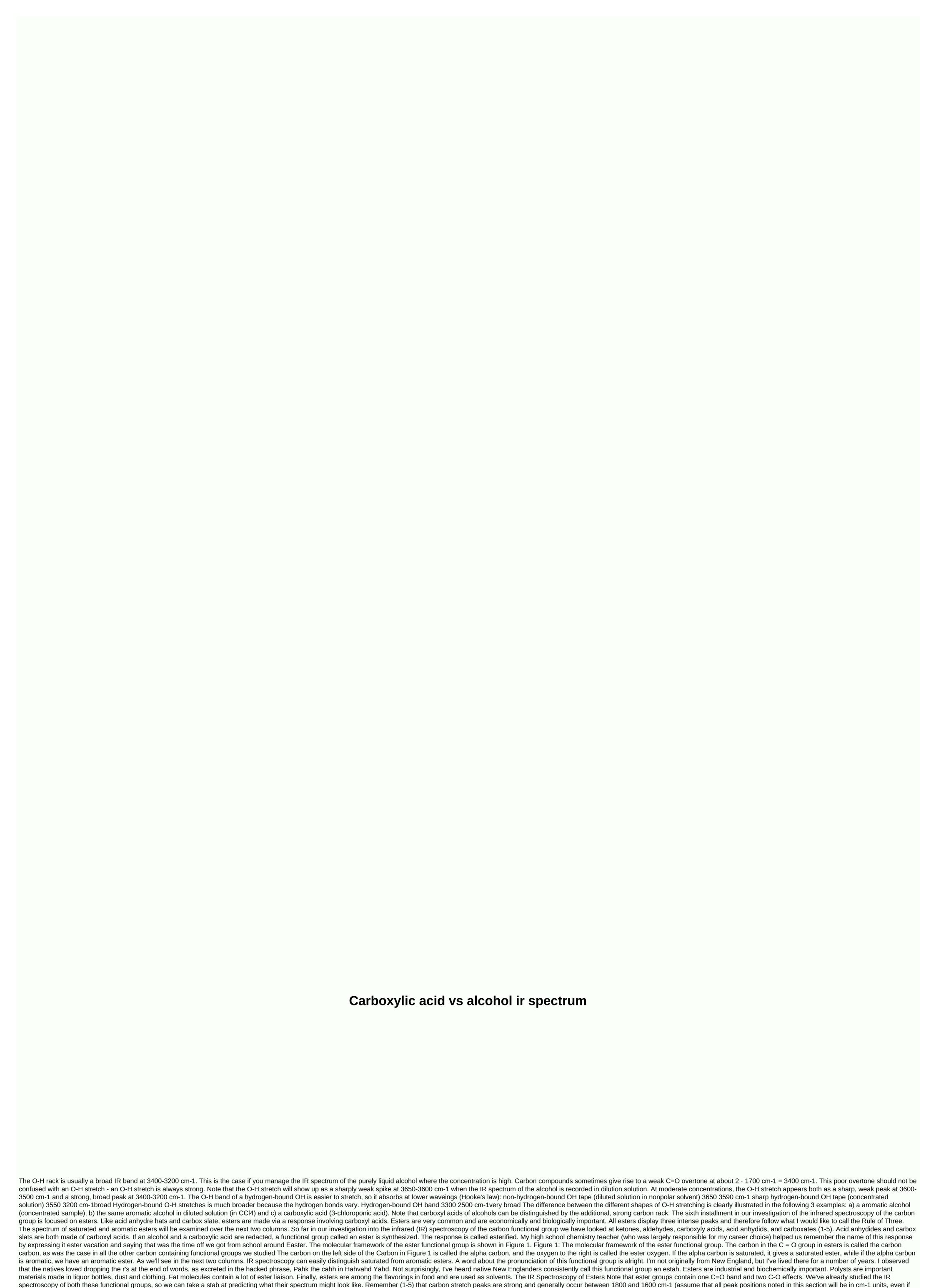
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not explicitly stated). We also know (6) that C-O stretches are intense peaks typically seen between 1300 and 1000. Since one of the C-O tyres in the ester group is attached to the carbon carbon and the other is not, we can expect the two to be chemically clear, have different power constant constants, and thus give rise to two separate peaks between 1300 and 1000. As it turns out, our predictions are correct. Esters have a memorable pattern of three intense peaks at ~1700, ~1200, and ~1100 of the C= O and two C-O stretches, and thus follow what I call the Rule of Three (7). An example of the

spectrum of a saturated ester, ethyl asetate, is shown in Figure 2. Figure 2: The IR spectrum of the saturated ethyl asetate is made of acetic acid and ethyl alcohol, one can in theory make it by reacting vinegar with vodka. Ethyl asetate becomes common in food and has a fruity flavor. This is a saturated ester because the alpha carbon is a methyl group. The Rule of Three peaks in Figure 2 is marked A, B and C and is easy to see, stitching like three long fingers in the middle of the spectrum. The peak at 1742, of course, is the carbon stretch, and for saturated esters in general this peak falls from 1755 to 1735. The second peak marked B on 1241 is off the ester oxygen, which is attached to the carbon carbon, and also the stretches of the alpha carbon-carbon C-C band. I call it the asymmetric C-C-O stretch and this vibration is illustrated in Figure 3. Figur the stretch of the second C-O band in the ester, which is the one on the right side of this tape, forming an O-C-C moiety. I call this vibration the asymmetric O-C-C stretch, and it is illustrated in Figure 4. Figure 4: The O-C-C rack of the ester functional group. This vibration accounts for the third of the Rule of Three peaks. For saturated esters in general, the O-C-C stretch appears from 1100–1030. To be clear, a linkage like C= Oh (O-CH3), which has an O-C tape rather than an O-C-C moiety, will still exhibit this peak. If you carefully reread this section, you might note that I lied to you. I said above that to saturated esters that the C-C-O stretch falls into the 1210-1160 series, and yet clearly the C-C-O piece of ethyl asetate falls at 1241. What's going on here? This spike is definitely a rule exception, but for once it's one we can understand and use of. Asetate esters are unique in that the alpha carbon is just a methyl group with nothing else attached. Remember from the first installment in this series (8) that one of the things that determines peak positions in IR spectroscopy is mass reduced, and that if the reduced mass of a functional group measured its IR peak positions in wavenumber. A methyl group is all about the lightest alpha carbon you can have in an ester, which is probably why asetate esters have a uniquely high wavenumber C-C-O stretch that typically falls around 1240. Acetate esters are common due to the ubiquity of aset acid. The fact that they have a unique asymmetric C-C-O stretching peak is useful in discriminating them from the many other types of saturated ester. Note in Figure 2 that as we go from left to right to peak A, B and C, the third peak is a little less intense than the other two. This intensity pattern is typical of esters and can be useful in identifying them. A summary of the group of wavenumbers for saturated esters is a common and economically important functional group made by reacting to an alcohol and a carboxylic acid. Their structural framework consists of a C=O group and two C-O effects. This gives rise to three intense peaks called the Rule of Three with peak positions at about 1700, 1200, and 1100 wavenumbers. The three vibrations is the C=O stretch, a C-C-O rack, and an O-C-C-C rack. Saturated esters were here, aromatic esters will be covered in the next installment. References (1) B.C. Smith, Spectroscopy 32(9), 31–36 (2017). (2) B.C. Smith, Spectroscopy 33(1), 14–20 (2018). (4) B.C. Smith, Spectroscopy 33(3), 16–20 (2018). (5) B.C. Smith, Spectroscopy 32(10), 28–34 (2017). (2) B.C. Smith, Spectroscopy 33(1), 14–20 (2018). (4) B.C. Smith, Spectroscopy 33(3), 16–20 (2018). (5) B.C. Smith, Spectroscopy 32(10), 28–34 (2017). (2) B.C. Smith, Spectroscopy 32(10), 28–34 (2017). (3) B.C. Smith, Spectroscopy 33(1), 14–20 (2018). 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